

# How to Renormalize the Schrödinger Equation III

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## Example: QED in Atoms — Ps, H, He...

## 8. QED Bound States

Top 10 Reasons  
to Work on  
Something Else

- Bound state = nonperturbative + multi-scale.

$$K \sim mv^2, P \sim mv, m \text{ where } v \approx \alpha$$

Eg) How big is  ?

$$= \alpha^3 m f(K/m, P/m)$$

$$= \alpha^3 m (\# + \# \alpha + \# \alpha^2 + \dots)$$

(=  $\alpha^5 m (\# + \dots)$  in Coulomb gauge!)

- QED renormalization  $\Rightarrow$  perturbation theory.  
(Feynman, Schwinger, Tomonaga...)

$\Rightarrow$  Bound state must be analyzed perturbatively:

$$\text{Eg)} \quad H = H_0 + \delta H_{\text{QED}} \quad \text{with} \quad H_0 = \frac{p^2}{2m} - \frac{\alpha}{r}$$

$$\Rightarrow E_0 = \frac{-\alpha^2 m}{2} \quad \psi_0 = \frac{e^{-r\alpha m}}{\sqrt{\pi}} \quad \delta H_{\text{QED}} = \text{Feyn. Diags.}$$

+ Rayleigh-Schrödinger Perturbation Th.

$\Rightarrow$  Big problem for He, Li... (no exact  $\psi_0, E_0!$ ).

- $\alpha$  expansion of  $E$ 's,  $\Gamma$ 's not so convergent.

Eg)

$$\Gamma_{\text{O-Ps}} = \Gamma (1 - 3\alpha + 4\alpha^2 - 20\alpha^3 + \dots)$$

$$\frac{3}{2\pi} \ln^2 \alpha + 0.7 \ln \alpha$$

$$\alpha^4 \ln^3 \alpha$$

→ Due to multi-scales:  
 $\ln^2(K/m) \dots$

QED renormalization

+

gauge non-invariance

+

relativity

+

nonperturbative bound  
state

⋮

⇒ TOO HARD!

Don't solve bound state  
problem directly in  
relativistic QFT.



# Effective Theory: Nonrelativistic QED System

- $p_e \sim m_e v \ll m_e$ 
  - $\Rightarrow$  Pair production highly suppressed.
  - $\Rightarrow$  e number doesn't fluctuate.
  - $\Rightarrow$  Don't need QFT for  $e$ .
- $p_\gamma \sim p_e \Rightarrow E_\gamma$  much larger:

$$(E_e \sim p_e^2/m_e \sim m_e v^2) \ll (E_\gamma \sim p_\gamma \sim m_e v)$$

$(\Rightarrow \text{photons} \approx \text{instantaneous})$

$\Rightarrow$  Remove photons from theory with energy cutoff


$$E_e \ll \Lambda_E < E_\gamma$$

$\Rightarrow$  Don't need QFT for  $\gamma$ .

- $\gamma$  interactions  $\rightarrow$  renormalization correction terms.
- Cutoff in  $E$  (not  $p$ )
  - $\Rightarrow$  Corrections local in  $t$  but **nonlocal in  $\mathbf{r}$** .
  - $\Rightarrow \gamma$  interaction  $\rightarrow$  instantaneous potentials  $V(\mathbf{r}), \sigma \cdot pV'(\mathbf{r}) \dots$

$\Rightarrow$  QED  $\rightarrow$  nonrelativistic Schrödinger theory:

$$H = \sum_i \frac{p_i^2}{2m_i} + V(\mathbf{r}_j, \mathbf{p}_j \dots) + \delta H(\text{v. soft } \gamma\text{s})$$


$p_\gamma \sim mv^2$  — couple weakly   
 $\Rightarrow E$ -dependent  $\delta V$  (eg, Lamb shift)

## Eg) Nonperturbative Positronium

Calculate O-Ps decay rate in two steps.

- 1) Define Hamiltonian ( $\equiv$  QED) with **finite** cutoff  $\Lambda$  built in.

$$H^{(\Lambda)} = \frac{\mathbf{p}^2}{m} - \frac{\mathbf{p}^4}{4m^3} + V + iW$$

Decay 

Define  $V, W$  by matching  $T = V + iW + V(E - H_0)^{-1}V + \dots$  to QED Feynman diagrams for  $e\bar{e}$  scattering, order-by-order in  $\alpha$  and  $p/\Lambda$ .

$$\langle \mathbf{l} | V | \mathbf{k} \rangle = -\frac{4\pi\alpha}{|\mathbf{k} - \mathbf{l}|^2} e^{-|\mathbf{k}-\mathbf{l}|^2/2\Lambda^2} \leftarrow \mathcal{O}(\alpha^2 m)$$

$$\mathcal{O}(\alpha^4 m) \rightarrow + \left[ \frac{\pi\alpha}{m^2} \frac{(l^2 - k^2)^2}{|\mathbf{k} - \mathbf{l}|^4} - \frac{4\pi\alpha}{2\Lambda^2} + \dots \right] e^{-|\mathbf{k}-\mathbf{l}|^2/2\Lambda^2}$$

$$\mathcal{O}(\alpha^5 m) \rightarrow + \frac{8\pi\alpha}{3\pi m^2} e^{-|\mathbf{k}-\mathbf{l}|^2/2\Lambda^2} \langle \mathbf{p} \cdot (H - E) \ln(\Lambda/(H - E)) \cdot \mathbf{p} \rangle + \dots$$

(Lamb Shift)

$$+ \frac{\alpha^2}{m^2} \left[ \frac{14}{3} \ln(|\mathbf{l} - \mathbf{k}|/m) - \frac{74}{15} - \frac{16}{3} \ln 2 + D \right] e^{-|\mathbf{k}-\mathbf{l}|^2/2\Lambda^2}$$

Match  $e\bar{e} \rightarrow e\bar{e}$

$$D = -\sqrt{\pi} \left[ \frac{-121}{36} \frac{\Lambda}{m} - 9 \frac{m}{\Lambda} + \frac{5}{3} \left( \frac{m}{\Lambda} \right)^2 \right] - \frac{16}{3} \ln \frac{\Lambda}{m}$$

Decay piece:

$$\langle \mathbf{l} | W | \mathbf{k} \rangle = \left[ A + B \frac{|\mathbf{k} - \mathbf{l}|^2}{m^2} \right] e^{-|\mathbf{k} - \mathbf{l}|^2 / 2\Lambda^2}$$

$A^{(0)}(1 + \alpha A^{(1)} + \dots)$   
from matching Born  
series for  $T$  with QED.

Not real QED behavior but  
equivalent for low-energy  
 $e$ 's, provided  $A, B$  correct.

$$\Rightarrow W(\mathbf{r}) \propto \left[ A - B \frac{\nabla^2}{m^2} \right] \delta_{1/\Lambda}^3(\mathbf{r})$$

$$A^{(1)} = a_0 + \frac{1}{\sqrt{\pi}} \left[ \frac{4}{3} \left( \frac{\Lambda}{m} \right) + 3 \left( \frac{\Lambda}{m} \right)^{-1} \right] + \mathcal{O} \left( \frac{\lambda}{m} \right),$$

$$\begin{aligned} A^{(2)} = & b_0 - 2a_1 + \frac{1}{\sqrt{\pi}} \left[ \frac{4}{3} \left( \frac{\Lambda}{m} \right) + 3 \left( \frac{\Lambda}{m} \right)^{-1} \right] a_0 + \frac{1}{3} \ln \frac{\Lambda}{m} \\ & + \frac{1}{\pi \sqrt{\pi}} \left\{ \left[ -\frac{44\sqrt{6}}{81} \left( \gamma - \ln \frac{2\Lambda^2}{3m^2} - 2 \right) \right] \left( \frac{\Lambda}{m} \right)^3 \right. \\ & + \left[ \frac{7}{3} \ln \frac{\Lambda}{m} + \frac{56\sqrt{6}}{27} \left( \gamma - \ln \frac{2\Lambda^2}{3m^2} - \frac{2}{7} \right) - \frac{37}{15} + \frac{1}{3} \ln 2 - \frac{7}{6} \gamma \right] \left( \frac{\Lambda}{m} \right) \left. \right\} \\ & + \left( \frac{83}{24\pi} - \frac{11\sqrt{3}}{12\pi} + \frac{11}{48} \right) \left( \frac{\Lambda}{m} \right)^2 \\ & + \left( \frac{25}{2\pi} - \frac{4\sqrt{3}}{3\pi} + \frac{17}{18} - \frac{5}{6} \ln 2 - \frac{1}{3} \gamma + \frac{2}{\sqrt{\pi}} \kappa \right) \\ & + \left( \frac{49}{6\pi} - \frac{3\sqrt{3}}{2\pi} - \frac{1}{4} \right) \left( \frac{\Lambda}{m} \right)^{-2} + \mathcal{O} \left( \frac{\lambda}{m} \right). \end{aligned}$$

## Note:

- Given  $V, W$  can solve theory even if completely ignorant of QED, renormalization, Effective Field Theory.
  - ◇  $\Lambda = m$  (or  $m/2$  or  $2m$ )  
⇒ No divergences ( $V$  analytic at  $r = 0$ ).
  - ◇ Renormalization built in, automatic.
  - ◇ High-order QED/relativity built in, automatic.
- Can solve **nonperturbatively** in  $V$  (eg, numerically).
  - ⇒ Don't need Rayleigh-Schrödinger perturbation theory.
  - ⇒ Move trivially to many- $e$  analysis (He...).
- No  $\ln \alpha$ s in  $V$  ⇒ perturbation theory for  $V$  is **more convergent** than for  $E_n$ .
  - ⇒ Compute  $V$  in (QED) perturbation theory; solve nonperturbatively in  $V$ .
  - ⇒ Schrödinger equation generates  $\ln \alpha$ s and resums them automatically.

2) Solve theory:

a) Diagonalize  $H^{((\Lambda))}$  (eg, on finite basis set of Gaussians).  
 $\Rightarrow E_n$ s and  $|\psi_n\rangle$ s.

b) Compute

$$\begin{aligned}\Gamma_n &= -2\langle\psi_n|W|\psi_n\rangle \\ &\quad + \bar{A}\langle\psi_n|e^{-r^2\Lambda^2/2}|\psi_n\rangle \\ &\quad - \bar{B}\langle\psi_n|\nabla^2 e^{-r^2\Lambda^2/2}|\psi_n\rangle \\ &\quad + \dots\end{aligned}$$

c) Publish numerical values obtained for  $\Gamma_n$ s.  
 $\Rightarrow (\Gamma_{1S} = 7.039967(10)\mu s^{-1}.)$



## QED in He

Traditional approach:

- Brute force numerical diagonalization of

$$H_0 = \frac{p_1^2}{2m} + \frac{p_2^2}{2m} - \frac{2\alpha}{r_1} - \frac{2\alpha}{r_2} + \frac{\alpha}{|r_1 - r_2|}$$

$\Rightarrow E_n^0$  to 14–17 digits;  $\psi_n^0$  to ? digits.

- Add QED + relativity using Rayleigh-Schrödinger P.Th. and numerical  $\psi_n^0$ s:

$$\delta E_n = \langle \psi_n^0 | \frac{-p_1^4}{8m^3} | \psi_n^0 \rangle + \dots$$

## Problems:

- Hard! Need expert in brute-force numerical analysis and QED.
- Numerical accuracy of  $\psi_n^0$  uncertain and well explored by highly singular QED/relativity corrections.

Solution:

- QED+relativity  $\rightarrow$  potential  $V_{\text{eff}}(r\dots)$   
(from Ps, H analyses).
  - Replace  $V_{\text{coul}}$  in brute-force evaluation by  $V_{\text{eff}}$ .
- $\Rightarrow$  Brute force numerics yields  $E_n$  that includes QED to 14-17 digits.  
( $\Rightarrow$  finished; no Step 2!)

## Numerical Analysis Bonus

- $-\alpha/r \rightarrow \infty$  as  $r \rightarrow 0$ .
  - $\Rightarrow$  Cusp in  $\psi(r)$  at  $r = 0$ .
  - $\Rightarrow$  Expansions (eg, on basis set) converge more slowly.
- Gaussian cutoff in  $V_{\text{eff}}(r) \Rightarrow$  analytic at  $r = 0$ .
  - $\Rightarrow \psi(r=0)$  analytic.
  - $\Rightarrow$  Expansions converge exponentially faster.

## Eg) Lamb Shift in H, He

- $N = \#$  basis functions  $< \infty$   
 $\Rightarrow$  effective cutoff  $\Lambda_N$ .
- Lamb shift

$$\psi^\dagger \text{---} \text{cloud} \text{---} \psi \sim \psi^\dagger \delta^3(r) \psi$$

$$\sim \left| \int^{\Lambda_N} d^3k \psi(k) \right|^2$$

where  $\psi(k) \rightarrow \frac{1}{k^4}$  for  $k$  large

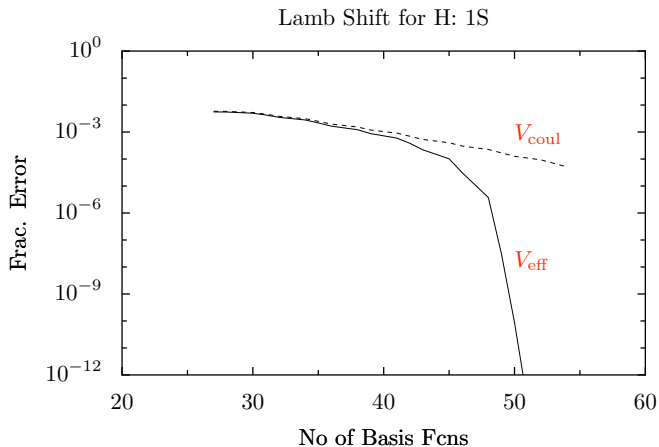
$$\Rightarrow \text{error} \propto 1/\Lambda_N^2$$

- $\psi_{\text{eff}} \sim e^{-k^2/2\Lambda^2}$  for  $k$  large

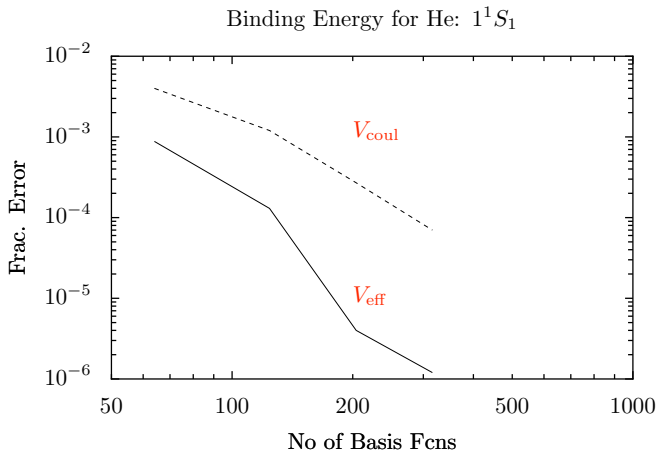
$$\Rightarrow \text{error} \propto e^{-\Lambda_N^2/\Lambda^2}$$

$\Rightarrow$  Errors exponentially suppressed.

# H(1S): Lamb Shift

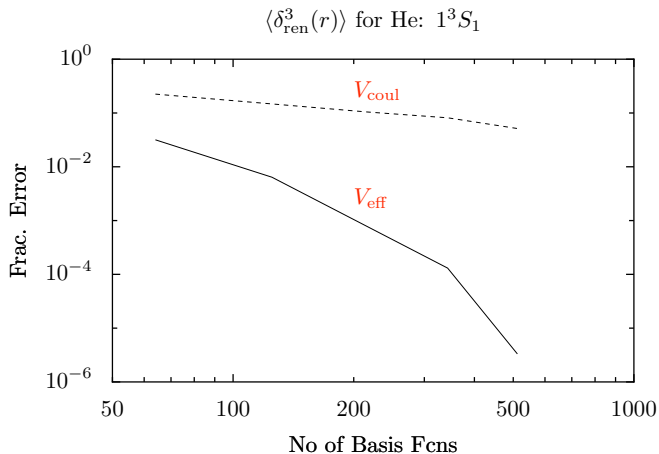


# He( $1^1S$ ): Binding Energy





# He( $1^1S$ ): QED/Relativistic Corrections



## Note:

- Computing  $\langle \psi_{\text{coul}} | \delta^3(r_1) | \psi_{\text{coul}} \rangle$  for He using  $\psi_{\text{eff}}$ .
- $\psi_{\text{eff}}(r_1 \approx 0, r_2)$  **totally different** from  $\psi_{\text{coul}}(r_1 \approx 0, r_2)$ .
- Renormalization  $\Rightarrow c, d$  such that

$$\langle \psi_{\text{coul}} | \delta^3(r_1) | \psi_{\text{coul}} \rangle = \langle \psi_{\text{eff}} | (c - d \nabla^2 / \Lambda^2) \delta_{1/\Lambda}^3(r_1) | \psi_{\text{eff}} \rangle$$

(N.B.  $c, d$  from H analysis — simple!)

# All-Orders Derivation: Short-Range Interactions

## 9. Short-Range Interactions

All interactions short-range  $\Rightarrow$  model by

$$\begin{aligned}\mathcal{L}^{(\Lambda)} = & \psi^\dagger \left( i\partial_t + \frac{\nabla^2}{2m} \right) \psi \\ & - \frac{g}{\Lambda m} (\psi^\dagger \psi)^2 - \frac{gd}{\Lambda^3 m} \{ \psi^\dagger \nabla^2 \psi, \psi^\dagger \psi \} - \frac{g_p}{\Lambda^3 m} \psi^\dagger \nabla \psi \cdot \psi^\dagger \nabla \psi - \dots \\ & - \frac{\hbar}{\Lambda^4 m} (\psi^\dagger \psi)^3 - \dots\end{aligned}$$

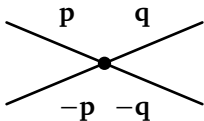
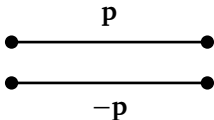
(Why not  $g/\Lambda^2$ ?)

For example:

- BEC atoms
- N-N interactions for  $p < 100$  MeV.
- ...

# Perturbation Theory: 2-Body

Leading order in  $1/\Lambda$  (c-of-m frame):



$$\frac{1}{E - \frac{p^2}{2m} - \frac{p^2}{2m} + i\epsilon} = \frac{-m}{p^2 + \gamma^2}$$

$$\gamma^2 \equiv -mE - i\epsilon$$

$$\frac{4\pi^2 g}{\Lambda m} \theta(p < \Lambda) \theta(q < \Lambda)$$

UV cutoff *separable!*

## All Orders in $g$

$$T(E) = \text{diagram}_1 + \text{diagram}_2 + \text{diagram}_3 + \dots$$

$$= \text{diagram}_1 \sum_{n=0}^{\infty} \left( \frac{\text{diagram}_2}{\text{diagram}_1} \right)^n$$

$$\Rightarrow T(E)^{-1} = \frac{1}{\text{diagram}_1} - \frac{\text{diagram}_2}{(\text{diagram}_1)^2}$$

$$- \frac{m}{4\pi^2} \int_0^{\Lambda} k^2 dk \frac{1}{k^2 + \gamma^2}$$

## IR vs UV

$$\int_0^\Lambda k^2 dk \frac{1}{k^2 + \gamma^2} = \int_0^\Lambda dk \frac{(k^2 + \gamma^2) - \gamma^2}{k^2 + \gamma^2} = \Lambda - \gamma^2 \int_0^\Lambda \frac{dk}{k^2 + \gamma^2}$$
$$= \Lambda - \gamma^2 \int_0^\infty \frac{dk}{k^2 + \gamma^2} + \gamma^2 \int_\Lambda^\infty \frac{dk}{k^2 + \gamma^2}$$

IR term  $\Rightarrow$

$$-\frac{\gamma\pi}{2}$$

$\Rightarrow$  odd power of  $\gamma$ .

UV term:  $\gamma \ll \Lambda \Rightarrow$

$$\gamma^2 \int_\Lambda^\infty \frac{dk}{k^2} \sum_{n=0}^{\infty} \left( \frac{-\gamma^2}{k^2} \right)^n$$
$$= \Lambda \left[ \frac{\gamma^2}{\Lambda^2} - \frac{\gamma^4}{3\Lambda^4} + \dots \right]$$

$\Rightarrow$  even powers of  $\gamma^2/\Lambda^2$ .

## Key Theorem (for later)

Prove by induction:

$$\int_0^\Lambda k^2 dk \frac{1}{k^2 + \gamma^2} \left( \frac{k^2}{\Lambda^2} \right)^n = \Lambda \left[ c_0 + c_1 \frac{\gamma^2}{\Lambda^2} + c_2 \frac{\gamma^4}{\Lambda^4} + \dots \right]$$

UV piece:  
⇒ even in  $\gamma$   
⇒ analytic in  $E$ .

$$- \frac{\gamma \pi}{2} \left( \frac{-\gamma^2}{\Lambda^2} \right)^n$$

IR piece:  $k^{2n} \rightarrow (-\gamma)^{2n}$   
⇒ odd in  $\gamma$   
⇒ non-analytic.



Final result (where  $\gamma = \sqrt{-mE}$ ):

$$\begin{aligned} T(E)^{-1} &= \frac{1}{\text{X}} - \frac{\text{O}}{(\text{X})^2} \\ &= \frac{\Lambda m}{4\pi^2 g} + \frac{m}{4\pi^2} \left[ \Lambda \left( 1 + \frac{\gamma^2}{\Lambda^2} + \dots \right) - \frac{\gamma \pi}{2} \right] \\ &= \frac{\Lambda m}{4\pi^2} \left( \frac{1}{g} + 1 \right) - \frac{m\gamma}{8\pi} + \mathcal{O} \left( \frac{\gamma^2}{\Lambda^2} \Lambda m \right) \end{aligned}$$

Remove cutoff dependence  
by tuning  $g \rightarrow g(\Lambda)$ ;  
 $\Rightarrow$  physical!

No coupling to tune away  
cutoff dependence;  
 $\Rightarrow$  unphysical *but suppressed*.

## Note:

- $T(E)^{-1} = (m/8\pi)(1/a - \gamma) + \dots$  where

$$\frac{1}{a} = \frac{2\Lambda}{\pi} \left( \frac{1}{g} + 1 \right)$$

defines physical “scattering length”  $a$ .

⇒ Cutoff independent physics with

$$g(\Lambda)^{-1} = \frac{\pi}{2a\Lambda} - 1$$

- $(1/g + 1) > 0 \Rightarrow$  single bound state with

$$\gamma_{\text{bd}} \equiv \sqrt{-mE_{\text{bd}}} = \frac{2\Lambda}{\pi} \left( \frac{1}{g} + 1 \right).$$

- Unphysical (and uninteresting) unless  $\gamma_{\text{bd}} \ll \Lambda$  (ie,  $g \approx -1$ ).  
( $\Rightarrow \mathcal{O}(\Lambda m(\gamma_{\text{bd}}^2/\Lambda^2))$  only suppressed by  $\gamma_{\text{bd}}/\Lambda$  relative to leading order.)
- “Natural” size of  $\gamma_{\text{bd}}$  is  $\mathcal{O}(\Lambda)$ ;  $\gamma_{\text{bd}} \ll \Lambda \Rightarrow$  tuning (BEC atoms) or luck (deuteron).

## All Orders in $1/\Lambda^n$ — $S$ -Waves

Include finite-size corrections to contact term:

$$\times \rightarrow \frac{4\pi^2 g}{\Lambda m} f(p^2/\Lambda^2) \theta(p < \Lambda) \theta(q < \Lambda) f(q^2/\Lambda^2)$$

$\Rightarrow$  still separable with

$$f(p^2/\Lambda^2) \equiv 1 + \sum_{n=1}^{\infty} d_n \left( \frac{p^2}{\Lambda^2} \right)^n$$

$\Rightarrow$  tuneable couplings:  $g, d_1, d_2 \dots$

## Note:

- Vertex  $p, q$  dependence:

$$f(p^2/\Lambda^2)f(q^2/\Lambda^2) = 1 + d_1 \left( \frac{p^2}{\Lambda^2} + \frac{q^2}{\Lambda^2} \right) + d_1^2 \left( \frac{p^2 q^2}{\Lambda^4} \right) + d_2 \left( \frac{p^4}{\Lambda^4} + \frac{q^4}{\Lambda^4} \right) + \dots$$

- But  $p^2 = q^2 = -\gamma^2$  on energy shell  $\Rightarrow p^2 q^2 \equiv p^4 \equiv q^4$  in correction terms for  $\mathcal{L}^{(\Lambda)}$ .

$$\Rightarrow (\psi^\dagger \nabla^4 \psi)(\psi^\dagger \psi) \equiv (\psi^\dagger \nabla^2 \psi)^2. \quad (\text{"Redundant" operator})$$

$\Rightarrow$  One coupling,  $d_n$ , enough to tune entire  $1/\Lambda^{2n}$  correction.

Final result (where  $\gamma = \sqrt{-mE}$ ):

$$T(E)^{-1} = \frac{1}{\text{diagram}} - \frac{\text{diagram}}{\left(\text{diagram}\right)^2}$$

$$= \frac{1}{f(p^2/\Lambda^2)f(q^2/\Lambda^2)} \left\{ \frac{\Lambda m}{4\pi^2 g} + \frac{m}{4\pi^2} \int_0^\Lambda k^2 dk \frac{f^2(k^2/\Lambda^2)}{k^2 + \gamma^2} \right\}$$

$$\frac{m}{4\pi^2} \left[ \Lambda \left( c_0 + c_1 \frac{\gamma^2}{\Lambda^2} + \dots \right) - \frac{m\gamma}{8\pi} f^2(-\gamma^2/\Lambda^2) \right]$$

$$\xrightarrow{q^2, k^2 \rightarrow -\gamma^2} \frac{\Lambda m}{4\pi^2} \left( \frac{1}{g} + t_0 + t_1 \frac{\gamma^2}{\Lambda^2} + t_2 \frac{\gamma^4}{\Lambda^4} + \dots \right) - \frac{m\gamma}{8\pi}$$

Renormalized:

$$T^{-1} = \frac{\Lambda m}{4\pi^2} \left( \frac{1}{g} + 1 + \frac{2d_1}{3} + d_1^2 + \dots \right)$$

$g$  removes  $\Lambda$  dependence.  
 $\Rightarrow$  Physical.  
N.B. Operator mixing.

$$+ \frac{\Lambda m}{4\pi^2} \left( \frac{\gamma^2}{\Lambda^2} \right) \left( d_1^2 - 2d_1 - \frac{1}{3} + \right. \\ \left. + 2d_1 \left( \frac{1}{g} + 1 + \frac{3d_1}{2} + \frac{d_1^2}{5} \right) + \dots \right)$$

$d_1$  removes  $\Lambda$  dependence.  
 $\Rightarrow$  Now physical!

$$+ \frac{\Lambda m}{4\pi^2} \left( \frac{\gamma^2}{\Lambda^2} \right)^2 \left( \dots \right) + \dots$$

$d_n$  fixes  $(\gamma^2/\Lambda^2)^n$  term.  
 $\Rightarrow$  Physical.

$$- \frac{m\gamma}{8\pi}$$

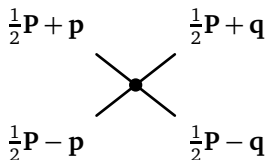
IR contribution  
 $\Rightarrow$  cutoff independent  
 $\Rightarrow$  physical.

## Note:

- Each  $d_n$  added and tuned  $\Rightarrow$  errors reduced by  $\gamma^2/\Lambda^2$ .
  - $\Rightarrow$  If leading  $\Lambda m$  term tuned/accidentally small, unlikely  $(\gamma^2/\Lambda^2)^n$  corrections have same suppression (although still relatively small).
  - $\Rightarrow$  Effective range ( $t_1$  term) can give more reliable indicator of scale of new physics than scattering length.
- No  $(\gamma^2/\Lambda^2)^n$  corrections to IR term  $-m\gamma/8\pi$ .
  - $\Rightarrow$  Good: no coupling constants to tune away  $\Lambda$  dependence.
- **Exercise:** Show there is still only one *physical* bound state.
- **Solution to all orders in  $\gamma^2/\Lambda^2$  and  $g$ !**
  - $\Rightarrow$  Universal behavior for *all* short-range potentials.
  - $\Rightarrow$  Effective range theory (Bethe, Schwinger, etc.).

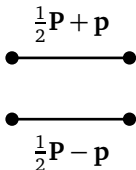


# Moving Frame: Galilean Invariance



$$\rightarrow \frac{4\pi^2 g}{\Lambda m} f(p^2/\Lambda^2) \theta(p < \Lambda) \theta(q < \Lambda) f(q^2/\Lambda^2)$$

Gal. Inv.  $\Rightarrow$  Independent of  $\mathbf{P}$ .



$$\rightarrow \frac{1}{E - \frac{(\mathbf{P}/2 + \mathbf{p})^2}{2m} - \frac{(\mathbf{P}/2 - \mathbf{p})^2}{2m}} = \frac{-m}{p^2 + \tilde{\gamma}^2}$$

$$\tilde{\gamma}^2 \equiv -m \left( E - \frac{p^2}{4m} \right)$$

$\Rightarrow T(E)^{-1}$  same but with  $\gamma \rightarrow \tilde{\gamma}$